

WEEKLY TEST TYM -1 TEST - 23 Rajpur Road SOLUTION Date 06-10-2019

[PHYSICS]

- 1. Kepler's second law is a consequence of conservation of angular momentum
- 2. According to Kepler's first law, every planet moves in an elliptical orbit with the sun situated at one of the foci of the ellipse.

In options (a) and (b) sun is not at a focus while in (c) the planet is not in orbit around the sun. Only (d) represents the possible orbit for a planet.

- 3. Kepler's law $T^2 \propto R^3$
- 4. During path *DAB* planet is nearer to sun as comparison with path *BCD*. So time taken in travelling *DAB* is less than that for *BCD* because velocity of planet will be more in region *DAB*.
- 5. Time period of a revolution of a planet,

$$T = \frac{2\pi r}{v} = \frac{2\pi r}{\sqrt{\frac{GM_S}{r}}} = \frac{2\pi r^{3/2}}{\sqrt{GM_S}}$$

- 6. Gravitational force is independent of the medium. Thus, gravitational force will be same i.e., F.
- 7. In arrangement 1, both forces act in the same direction. In arrangement 3, both the forces act in opposite direction. This alone decides in favour of option (a),
- 8. If a point mass is placed inside a uniform spherical shell, the gravitational force on the point mass is zero. Hence, the gravitational force exerted by the shell on the point mass is zero.

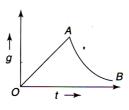
9.
$$g_d = g\left(1 - \frac{d}{g}\right)$$

or $g_d = g\frac{R - d}{R}$



or
$$g_d = \frac{gy}{R}$$
 or $g_d \propto y$

So, within the Earth, the acceleration due to gravity varies linearly, with the distance from the centre of the Earth. This explains the linear portion *OA* of the graphs.



10. The value of g at the height h from the surface of earth

$$g' = g\left(1 - \frac{2h}{R}\right)$$

The value of g at depth x below the surface of earth

$$g' = g\left(1 - \frac{x}{R}\right)$$

These two are given equal, hence $\left(1 - \frac{2h}{R}\right) = \left(1 - \frac{x}{R}\right)$ On solving, we get x = 2h

11. Acceleration due to gravity $g = \frac{4}{3}\pi\rho GR$ $\therefore g \propto \rho R$ or $\frac{g_m}{g_e} = \frac{\rho_m}{\rho_e} \cdot \frac{R_m}{R_e}$

As
$$\frac{g_m}{g_e} = \frac{1}{6}$$
 and $\frac{\rho_e}{\rho_m} = \frac{5}{3}$ (given)

$$\therefore \frac{R_m}{R_e} = \left(\frac{g_m}{g_e}\right) \left(\frac{\rho_e}{\rho_m}\right) = \frac{1}{6} \times \frac{5}{3} \therefore R_m = \frac{5}{18} R_e$$

12. Acceleration due to gravity $g = \frac{GM}{R^2}$

$$\therefore \frac{g_{\text{moon}}}{g_{\text{earth}}} = \frac{M_{\text{moon}}}{M_{\text{earth}}} \cdot \frac{R_{\text{earth}}^2}{R_{\text{moon}}^2} = \left(\frac{1}{80}\right) \left(\frac{4}{1}\right)^2$$
$$g_{\text{moon}} = g_{\text{earth}} \times \frac{16}{80} = \frac{g}{5}$$

13. Acceleration due to gravity $g = \frac{4}{3}\pi\rho GR$

$$\therefore g_1: g_2 = R_1 \rho_1: R_2 \rho_2$$

14.
$$g' = g\left(1 - \frac{d}{R}\right) \Rightarrow \frac{g}{4} = g\left(1 - \frac{d}{R}\right) \Rightarrow d = \frac{3R}{4}$$

15. We know
$$g = \frac{GM}{R^2} = \frac{GM}{(D/2)^2} = \frac{4GM}{D^2}$$

If mass of the planet = M_0 and diameter of the planet = D_0 . Then $g = \frac{4GM_0}{D_0^2}$.